

Planetary Overview

Courtesy: NASA



Inventory

- ☼ 1 Sun, containing 99.9% of mass
- ☼ 8 planets; 5 dwarf planets, and counting
- ☼ 15 moons over 1000 km in diameter
 - ∩ many smaller moons
- ☼ ~100,000 asteroids
- ☼ probably billions of Kuiper belt objects, scattered disc objects and Oort cloud objects
 - ∩ including billions of potential comets
- ☼ billions of meteorites, meteoroids and debris
- ☼ solar wind
- ☼ magnetic field



Bode's Law of Planetary Distances

☼ Empirical relationship (discovered by Titius!),
with no basis yet in theory

☼ If a is the average
distance of a planet
to the Sun, then:

$$a = \frac{3n + 4}{10} \text{AU}, \quad n = 0, 1, 2, 4, 8, 16, 32$$

$n = 0$ for *Mercury*

$n = 1$ for *Venus*

$n = 2$ for *Earth*

$n = 4$ for *Mars*

$n = 8$ for? – Ceres (1801)

$n = 16$ for *Jupiter*

$n = 32$ for *Saturn*

$n = 64$ for? – Uranus (1784)



How the Planets fit Bode's Law

★ After
K & K
chapter
7 table

n	Bode's Law Prediction	Today's Measured Distance (AU)	Object
0	0.4	0.39	Mercury
1	0.7	0.72	Venus
2	1	1	Earth
4	1.6	1.52	Mars
8	2.8	2.8	Ceres
16	5.2	5.2	Jupiter
32	10	9.54	Saturn
64	19.6	19.19	Uranus
		30.06	Neptune
128	38.8	39.4	Pluto



Masses and Densities

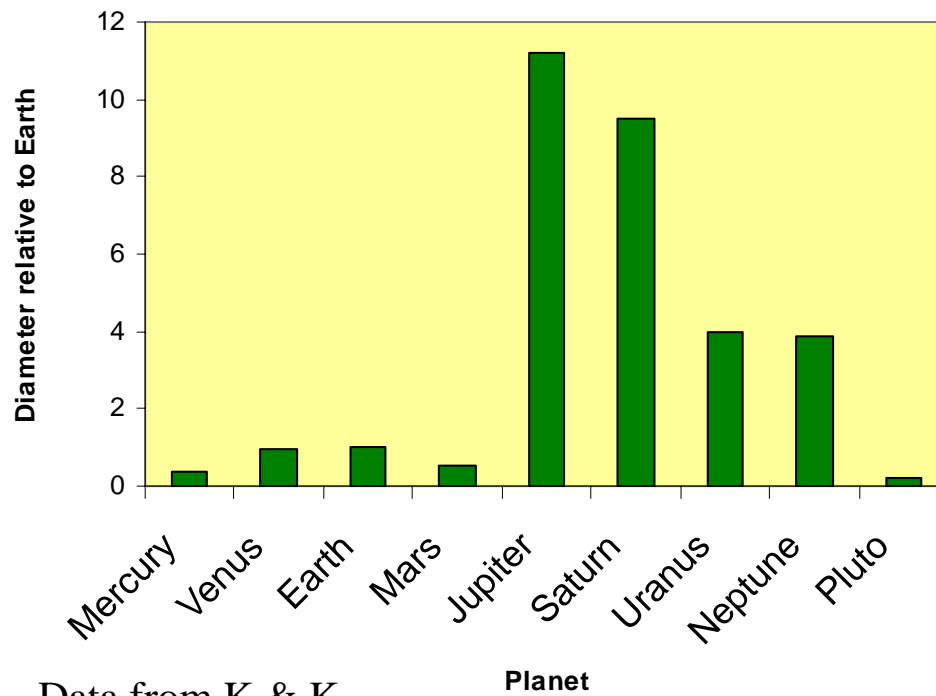
- ☼ Masses of all planets with moons are obtained through Kepler's 3rd Law
- ☼ Densities are calculated from ratio

$$\text{density (kg m}^{-3}\text{)} = \frac{\text{mass(kg)}}{\text{volume(m}^3\text{)}}$$

- ☼ **Terrestrial** planets are mainly rocky matter, with a molten core
- ☼ **Jovian** planets are mainly gas and liquid, with a rock core

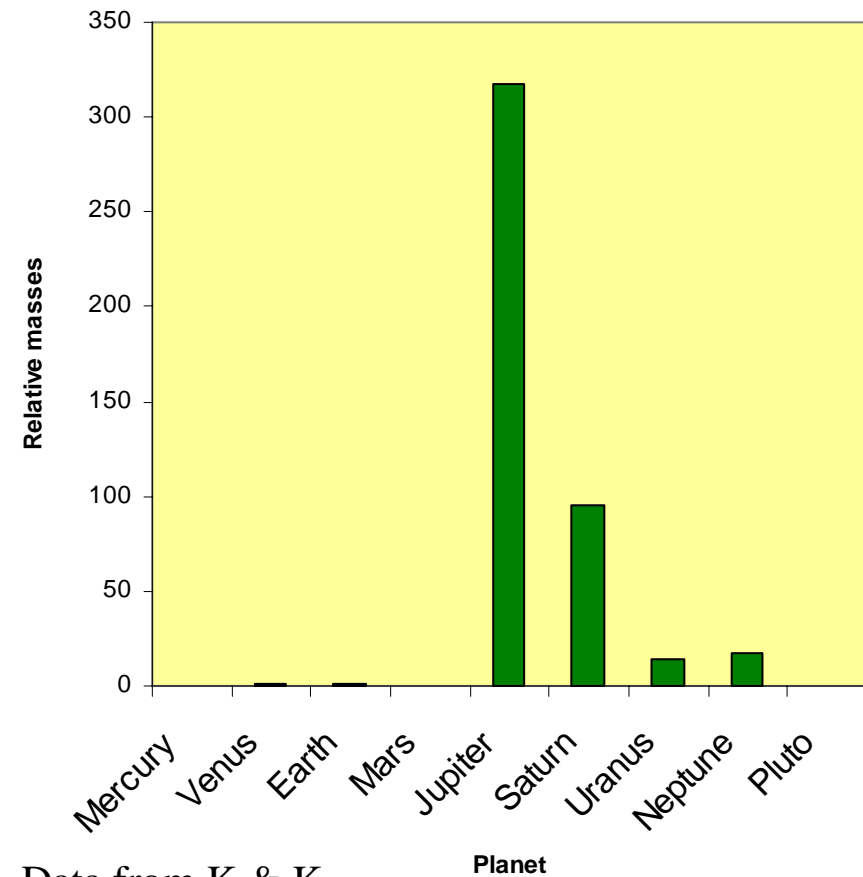
Diameters & Masses

Planetary diameters relative to Earth



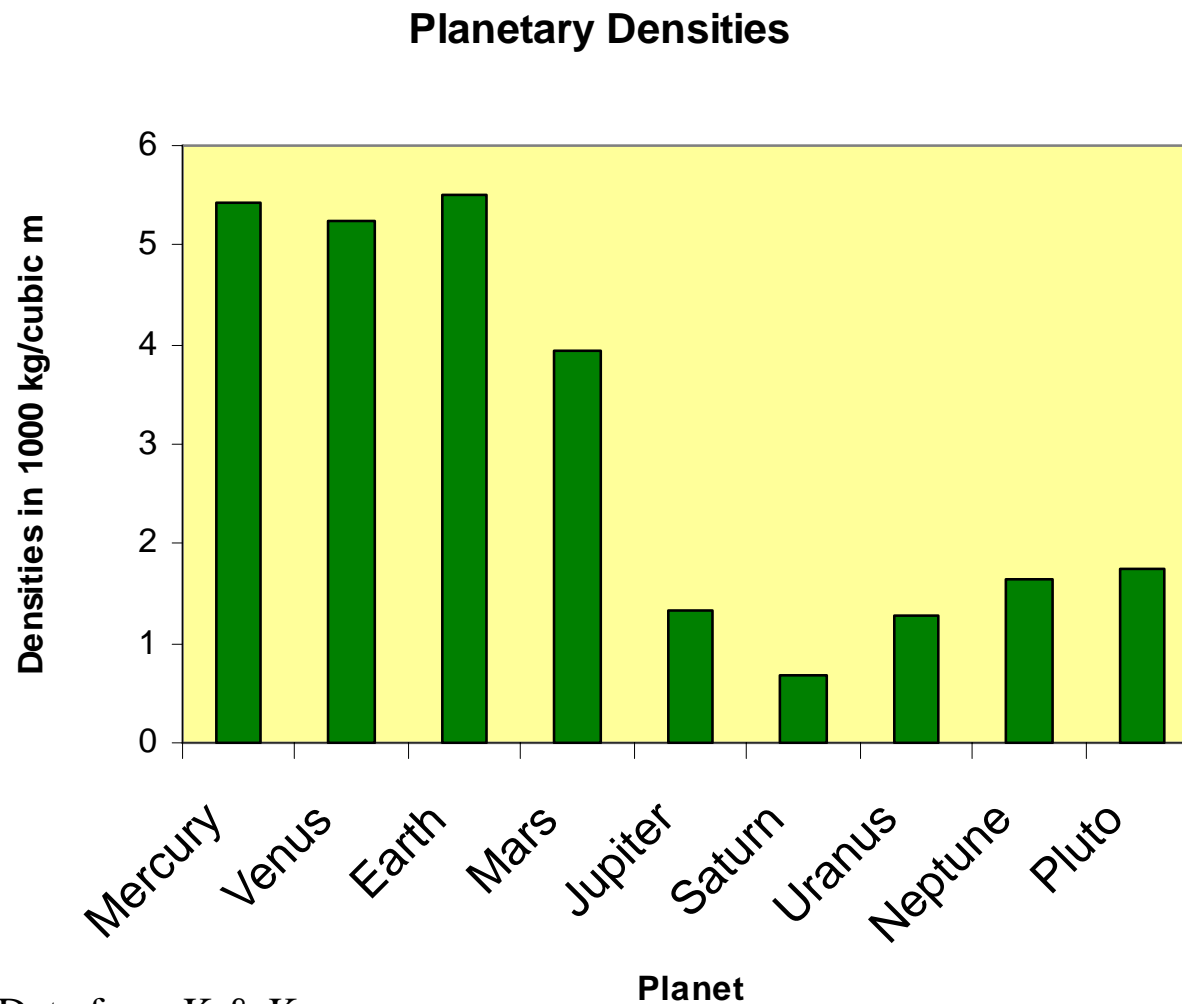
Data from K & K

Planetary masses relative to Earth



Data from K & K

Planetary Densities



Data from K & K



Orbits and Satellites

- ★ Rotation is generally in the same direction:
 - ★ rotation of Sun (equator rotates faster than polar regions)
 - ★ rotation of planets about their axes (except Venus, Uranus and Pluto)
 - ★ orbits of planets about Sun
 - ★ orbits of moons about planets
 - ♊ >100 named moons
- ★ Most orbits nearly circular and in same plane
- ★ All Jovian planets have ring systems



Atmospheres

★ Average energy ($\frac{1}{2}mv^2$) of molecules \propto temp

★ lighter molecules therefore move faster

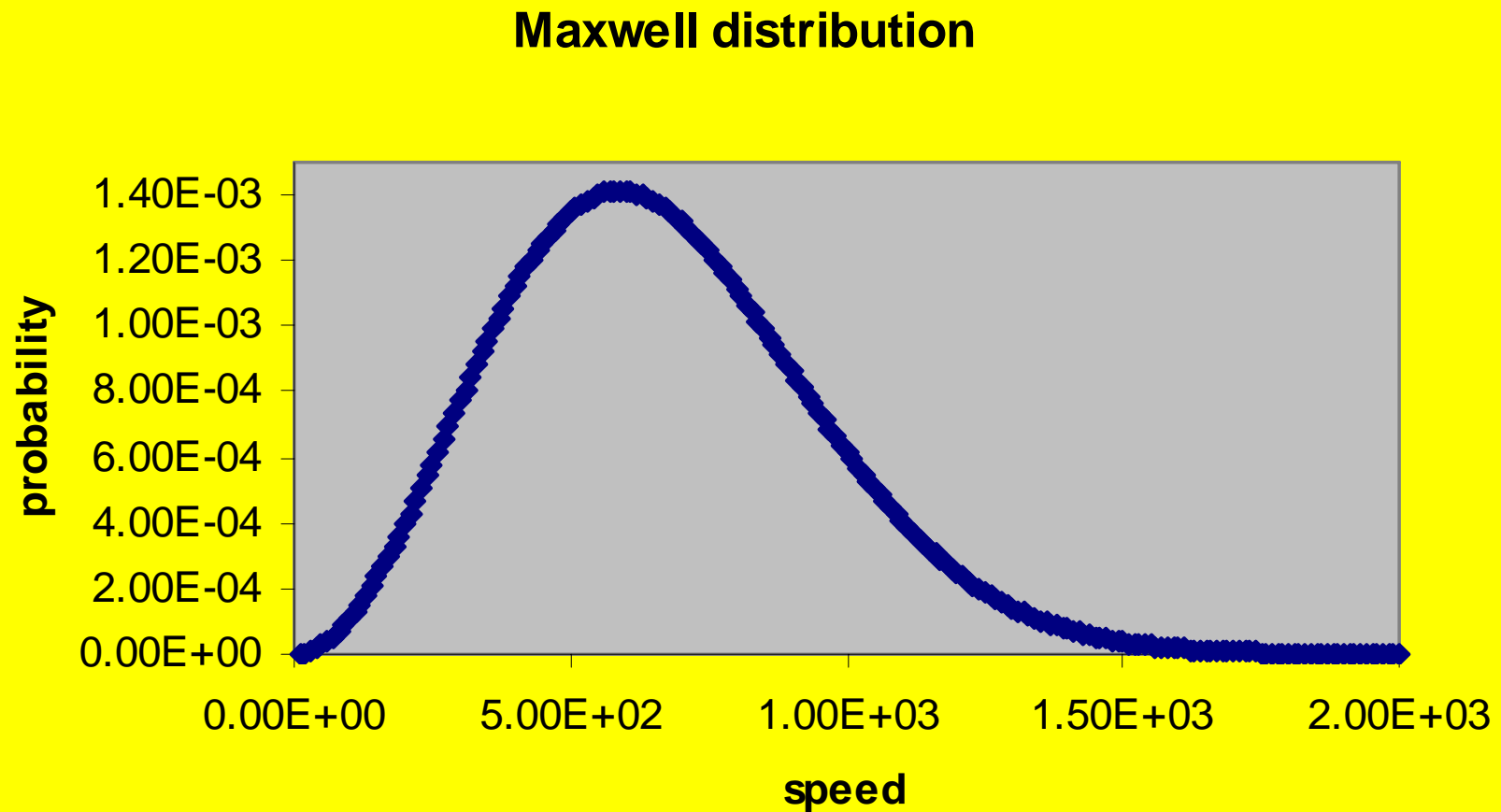
★ In a gas there is a wide range of molecular speeds, called the Maxwell distribution → [next slide]

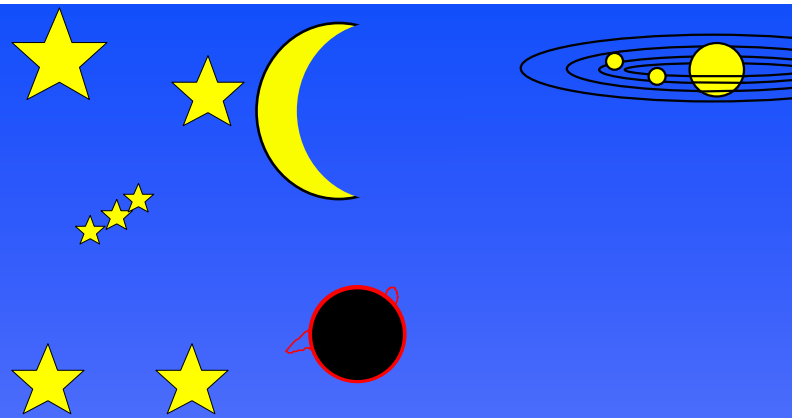
★ a significant number of molecules travel at more than 10 times the average speed

★ if they reach the escape velocity without colliding any more, they will escape

★ The Earth has lost its primitive H_2 and He

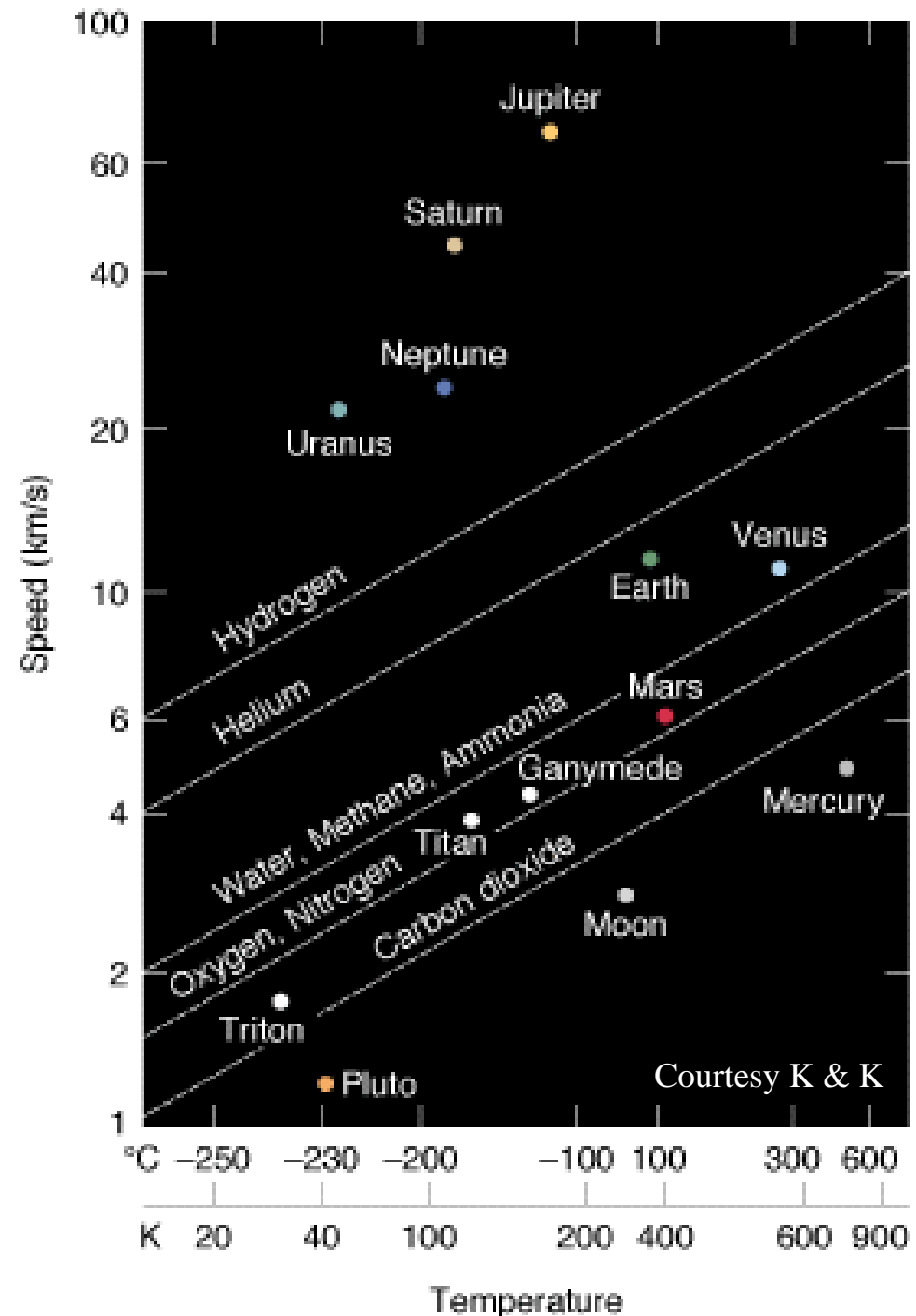
Maxwell Distribution





How Molecular Speeds Depend on Temperature

Each planet is plotted at its
atmospheric temperature
and escape velocity





Features of the Solar System

- ☼ system nearly planar
 - ☼ angular momentum mostly in outer regions
 - ☼ spacing of planets increases with distance from Sun
 - ☼ chemical composition of planets
 - ☼ cratering everywhere
 - ☼ ring systems on Jovian planets
 - ☼ presence of asteroids, comets & meteorites
- ★ Planetary systems likely to be common around other single stars — prediction prior to planet discoveries

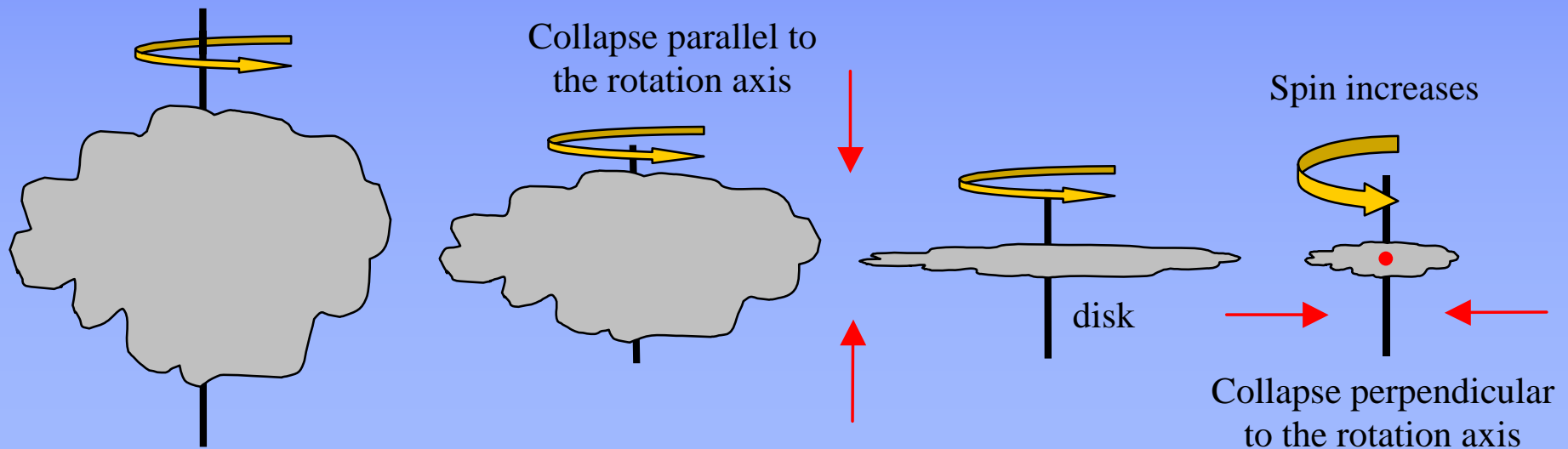


Formation of Solar System

- ★ **Catastrophe** theories consider the Solar system pulled out from a star
 - ★ the physics and chemistry of catastrophe theories cannot be made to produce the observed features
- ★ **Evolutionary** theories describe formation from an initial large cloud of rotating gas
 - ★ gravitational attraction along with conservation of angular momentum gives the condensing cloud a disk shape

Collapse of a rotating gas cloud

- ★ Collapse parallel to the rotation axis does not redistribute angular momentum
- ★ Collapse perpendicular to the rotation axis causes the cloud to spin faster





Modern Evolutionary Theory

★ von Weizäcker's analysis of a rotating gas condensing around a *protosun* showed that the gas would form eddies, with larger eddies further from the centre

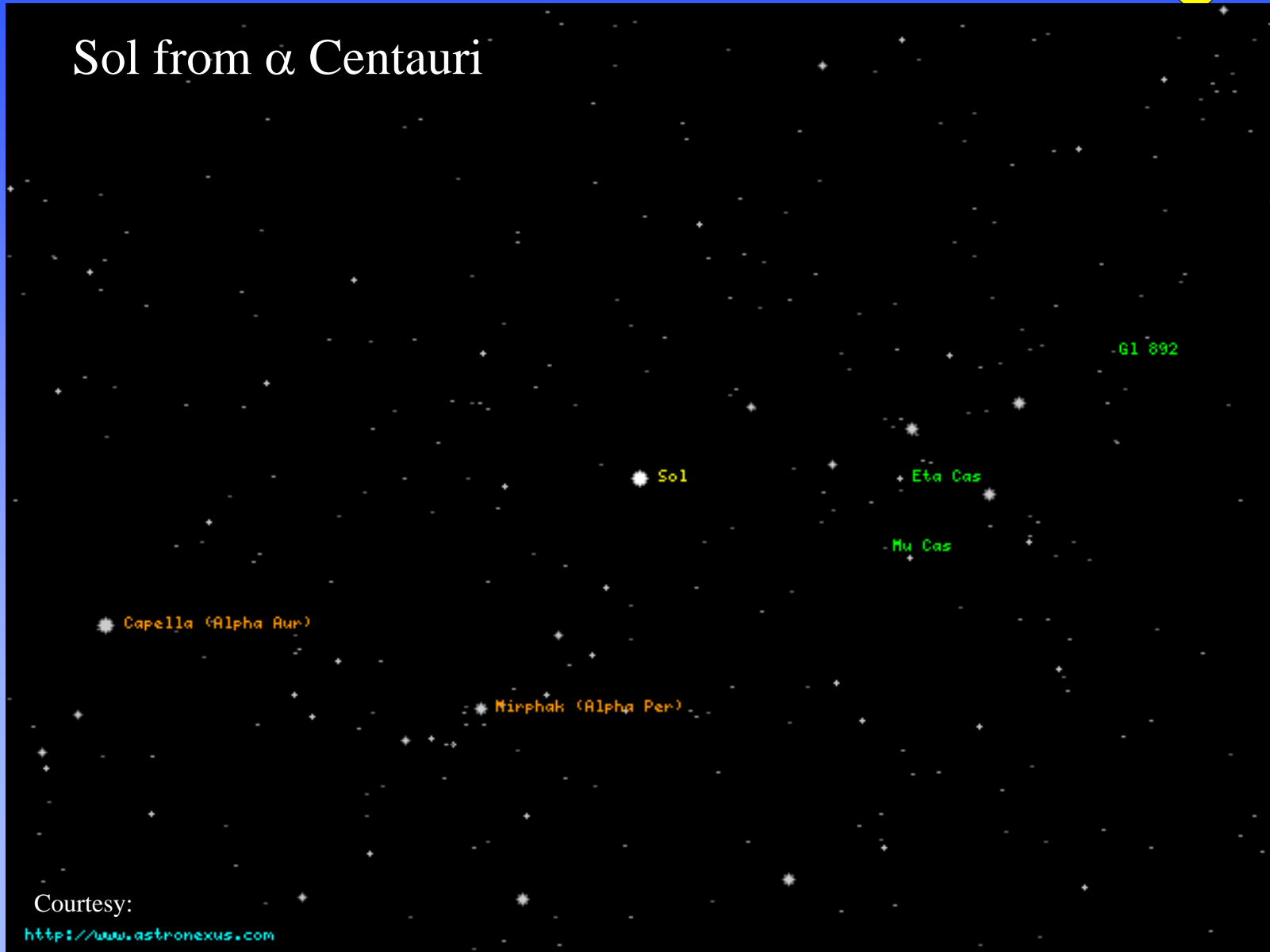
★ *Planetessimals* slowly formed by collision

★ asteroids are remnant planetesimals that failed to join together due to the stirring influence of Jupiter

★ The expected fast rotation of the Sun was slowed by the influence of its magnetic field on the ionised gas it created

Planets around stars

Sol from α Centauri

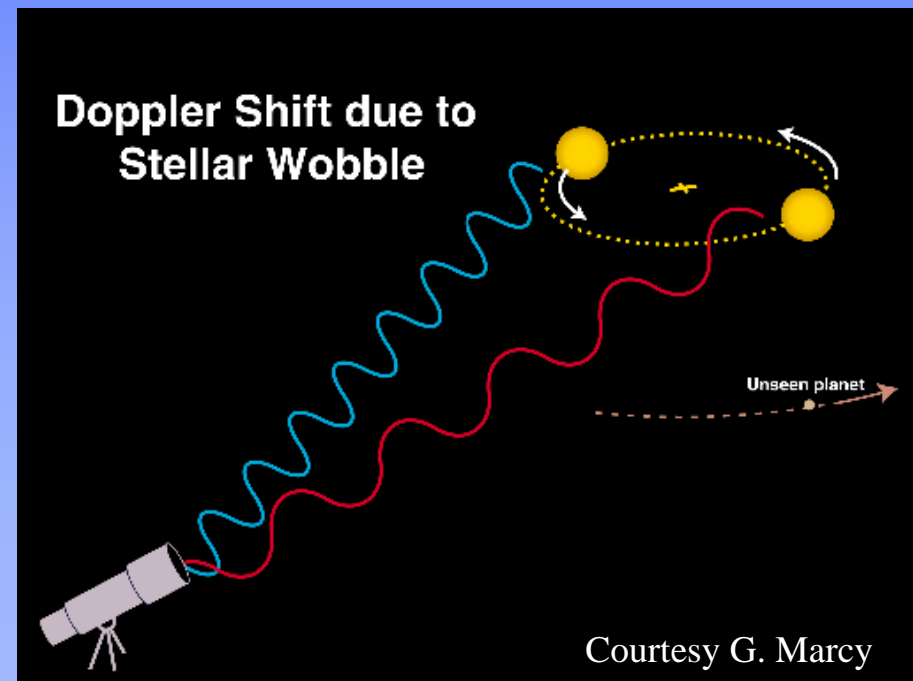


Courtesy:

<http://www.astronexus.com>

Extrasolar Planet Search

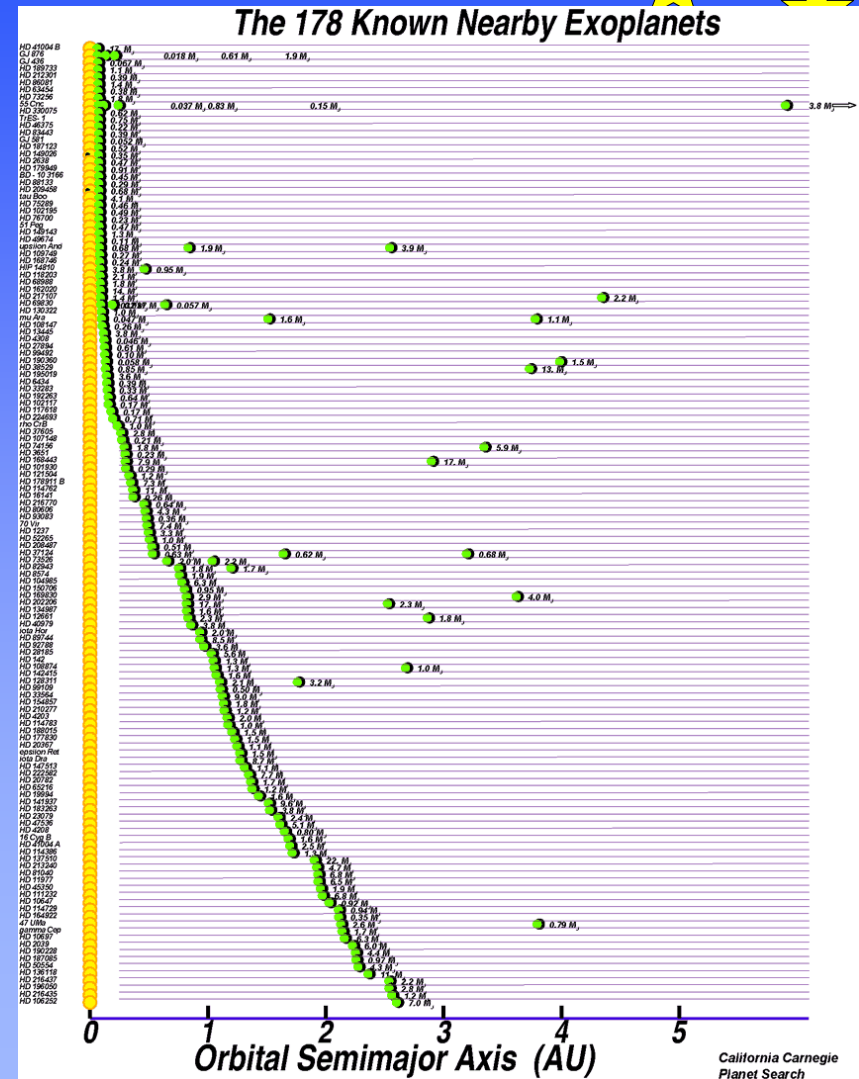
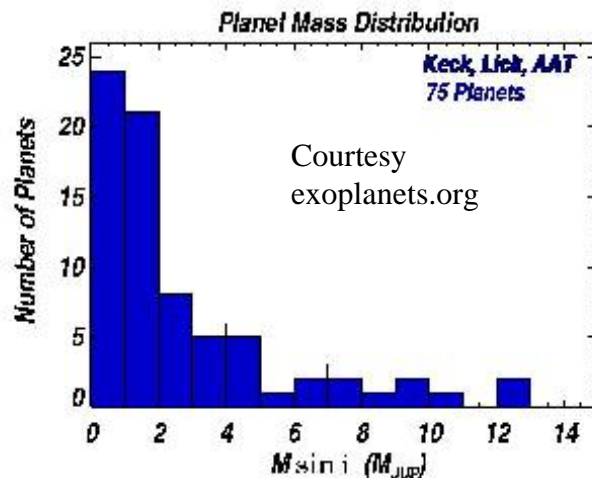
- ★ Some 300 stars now have confirmed planets
- ★ All candidates within ~ 100 LY of Earth have been examined
- ★ Most common technique is to detect very small Doppler shifts of spectral lines from parent stars



Extrasolar Planet Properties

★ Most planets detected are

- ★ close to their stars
- ★ have masses like Jupiter
- ★ are in eccentric orbits

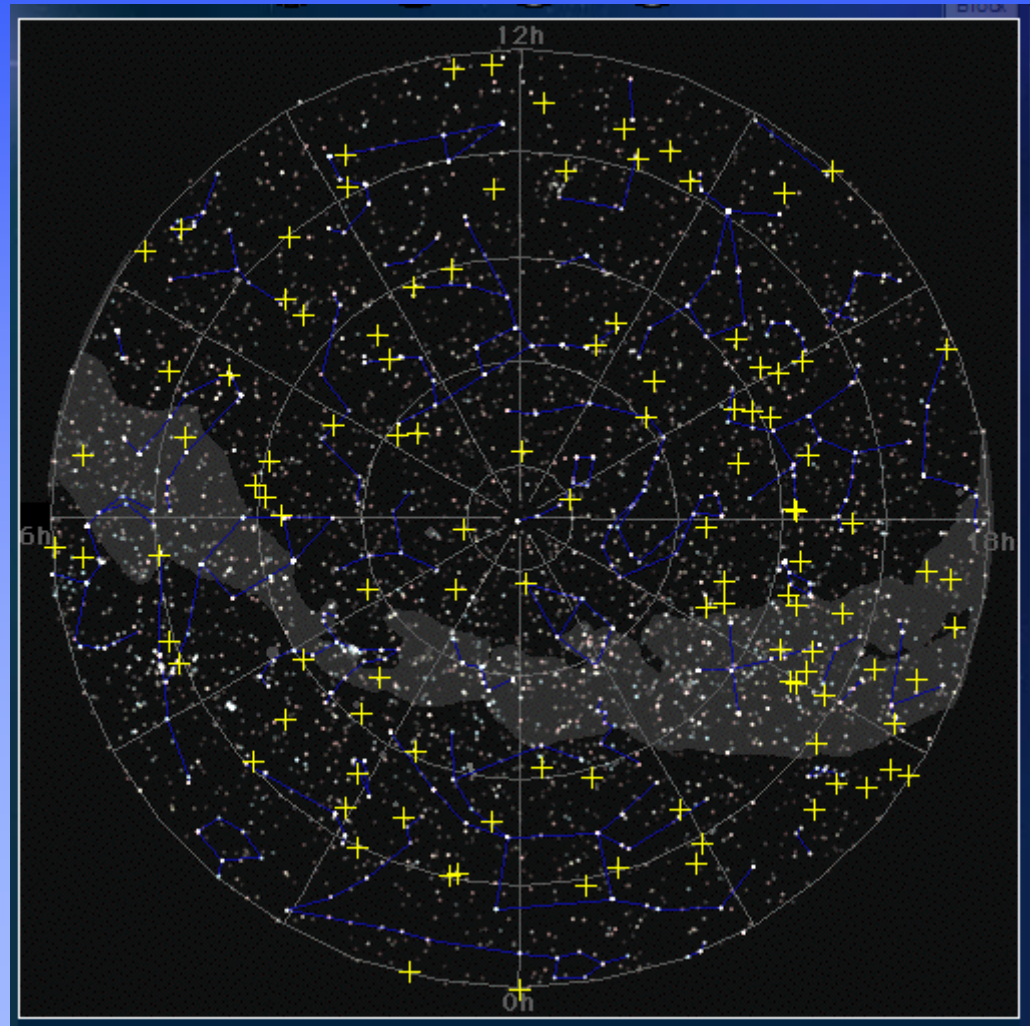


Courtesy: <http://exoplanets.org/massradiiframe.html>

Where are they?

- ★ Everywhere!
- ★ N celestial hemisphere →
- ★ For 3D skymap

☾ <http://media4.obspm.fr/exoplanets/base/carte3d.php>



Courtesy: <http://media4.obspm.fr/exoplanets/base/carte.php>

A Shift in Expectations

- ★ Large planets provide the biggest wobble of their parent stars

- ★ the closest large planet in the solar system (Jupiter) orbits in ~12 years

- ★ looking for changes in stars over times as long as this requires patience and instrument stability

- ★ The new perception is the discovery that large planets that circle their parent in only a few days exist around some stars

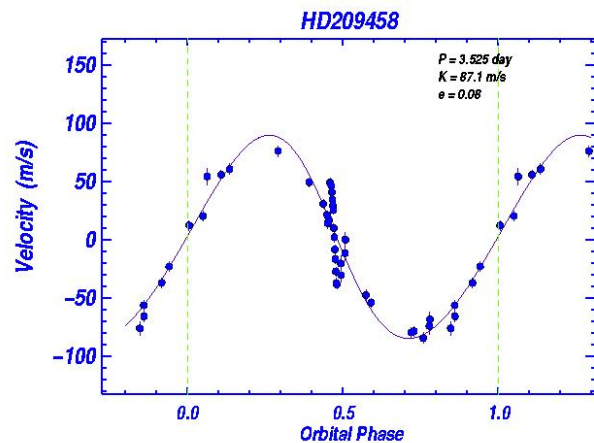
- ★ planetary systems like these are unlike the solar system



Example of HD209458

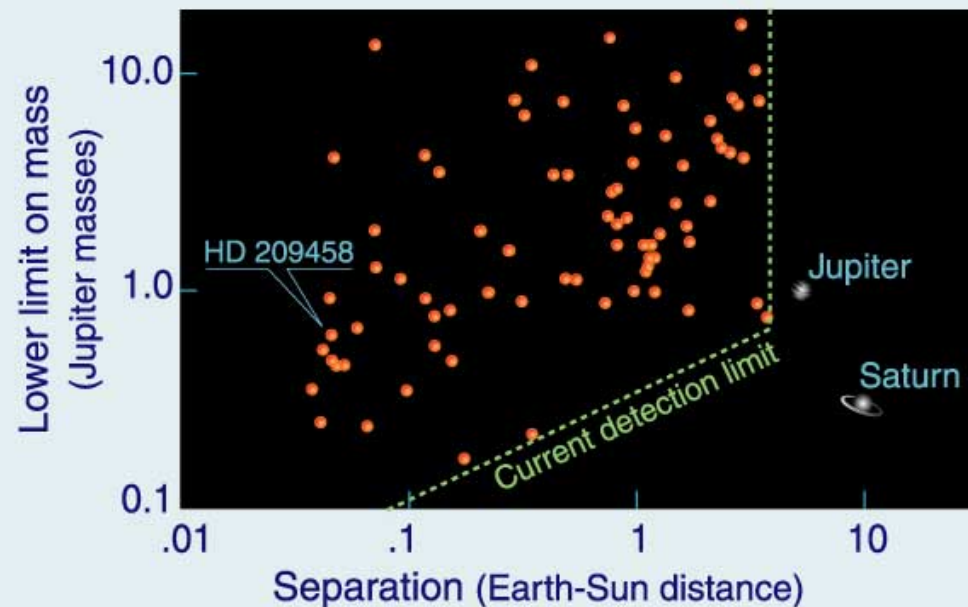
★ HD209458 is a Sun-like star 150 LY distant in the constellation of Pegasus

★ it has a planet $0.7M_{\text{Jup}}$ orbiting in 3.5 days



Courtesy: Geoff Marcy ↑

Discovery space for extrasolar planets

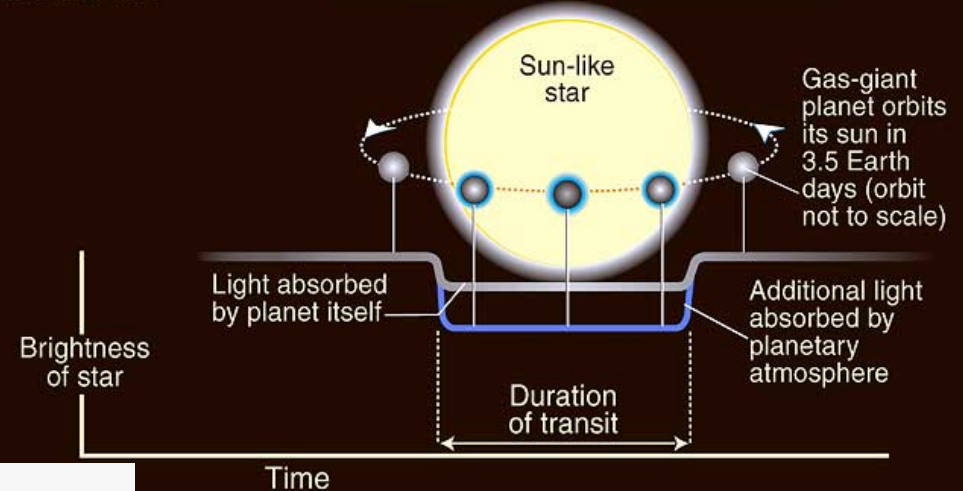
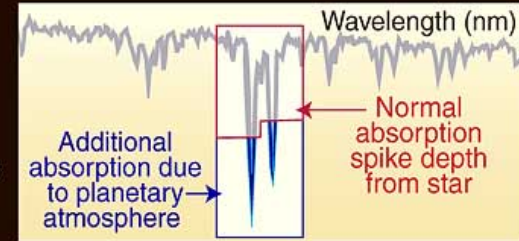


Courtesy: A. Feild ↑

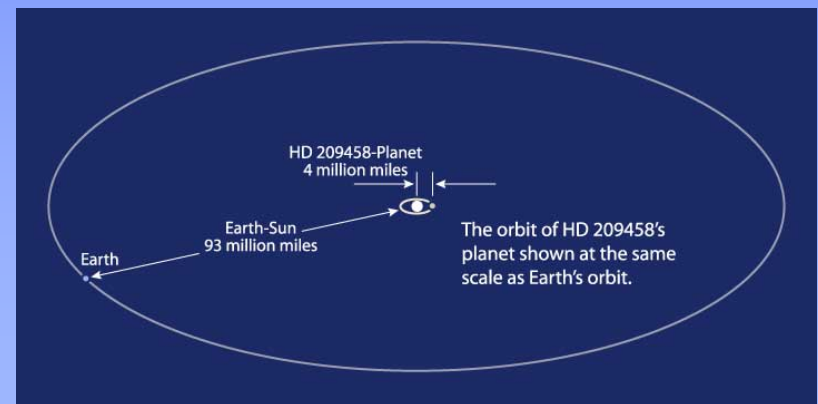
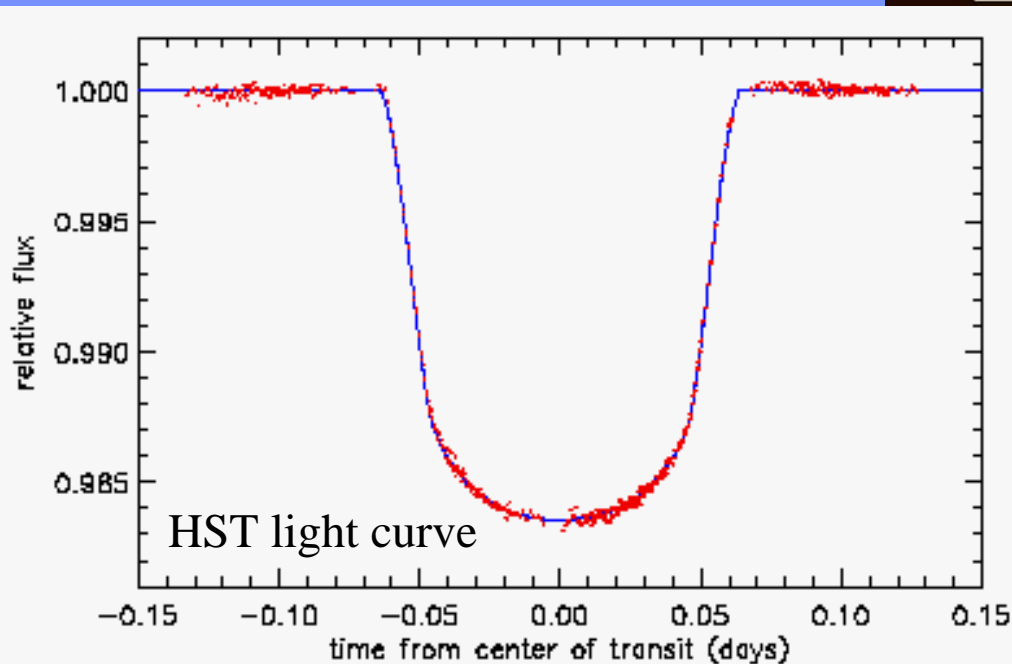
Transit of HD209458

- ★ The planet transits the star, affecting the light received
- ★ The planet is bigger than Jupiter

HST detects additional sodium absorption due to light passing through planetary atmosphere as planet transits across star



Courtesy: A. Feild ↑

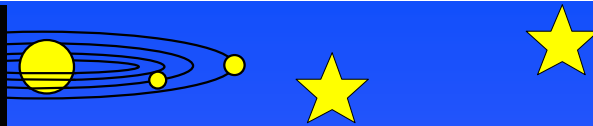


Courtesy: Z. Levay ↑

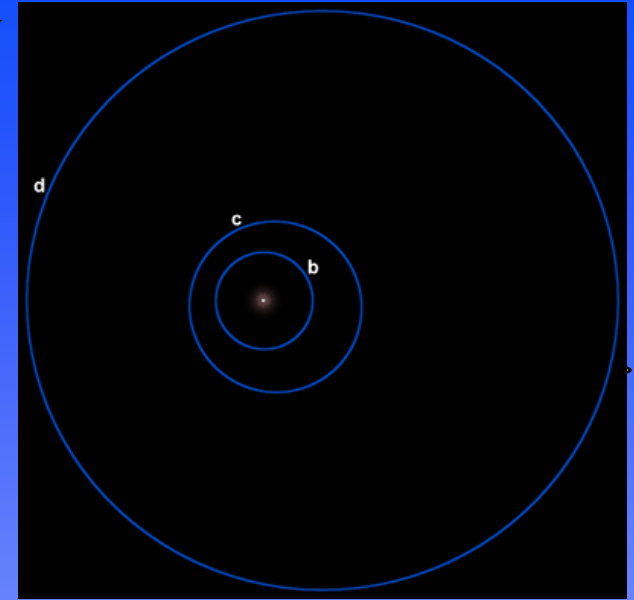


Looking for Earth-like Planets

- ★ Giant planets close in to their parent stars will have temperatures of over 1000°C facing the star
- ★ Detecting Earth-like planets will not be easy
 - ☼ 47 UMa has at least 2 giant planets in circular orbit at a distance of several AU
 - ☼ a new era in astronomy has dawned
 - ☼ new instruments, including giant mirrors and custom designed space probes
 - ☼ new techniques such as looking for the dip in light as a planet transits its parent star

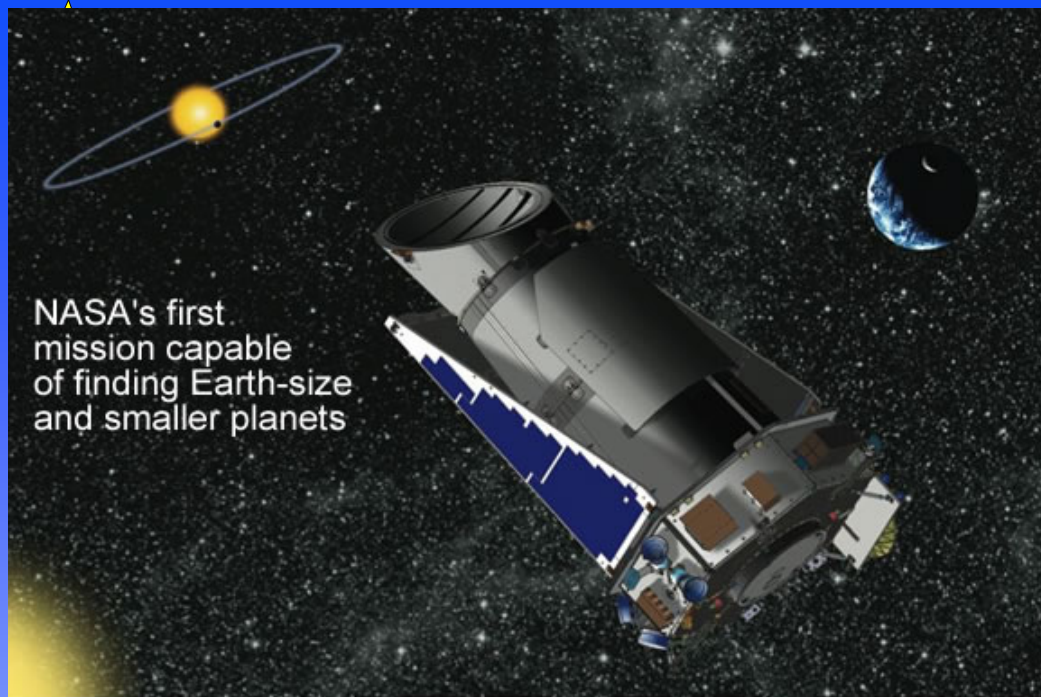


Gliese 581



- ★ Nearby star: 20 LY distance
- ★ Red dwarf star; $m = 10.5$; $T \sim 3500$ K; $\sim 1\%$ output of Sun; mass $\sim 0.3M_{\odot}$
- ★ 3 planets discovered by wobble technique
- ★ Outer 2 (Gliese 581 c & Gliese 581 d) discovered in 2007 at either edge of habitable zone; both larger than Earth
- ★ May be most Earth-like planets found; may be more like large Venus and large Mars, too hot and too cold

Kepler Mission



- ★ Mission to look for Earth-like planets using the transit dimming technique
- ★ Monitoring light emission from 100,000+ stars to a precision of 20 ppm in a fixed area of sky for 4 – 6 years
- ★ Launch Feb 2009 into Earth-trailing heliocentric orbit

