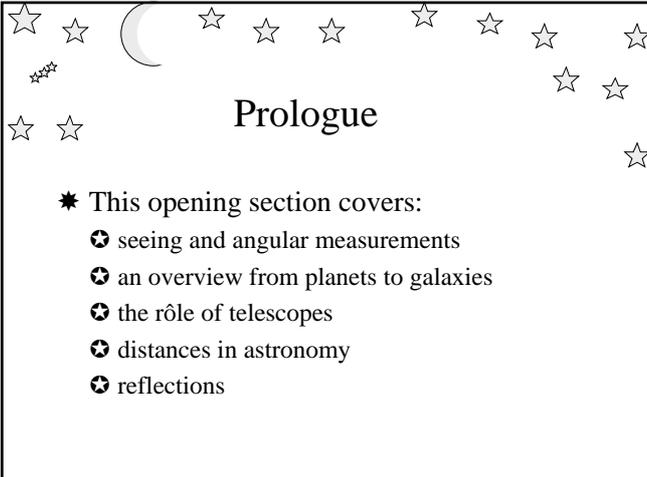
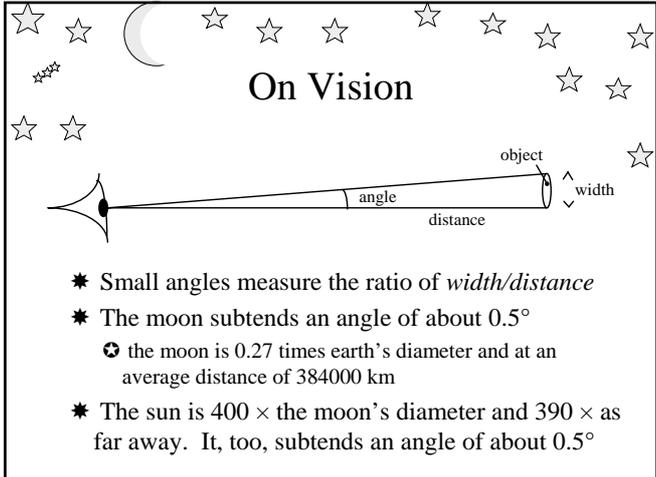
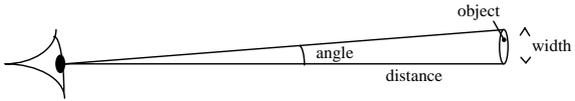

  
**Astronomy**
  
 Department of Physics
   
 University of Aberdeen
   
  
 A level 1 course
   
 by
   
**John S. Reid, FRAS**



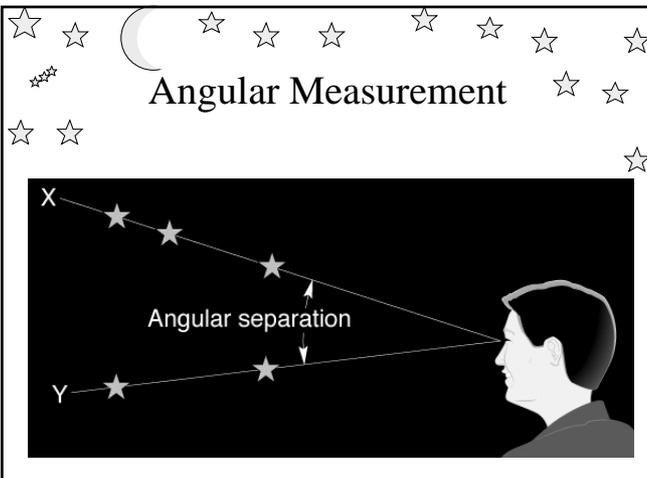

  
**Prologue**

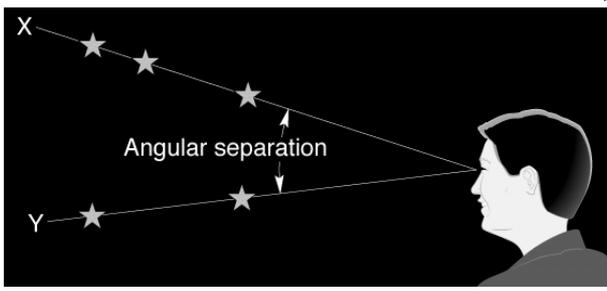
- \* This opening section covers:
  - ⊛ seeing and angular measurements
  - ⊛ an overview from planets to galaxies
  - ⊛ the rôle of telescopes
  - ⊛ distances in astronomy
  - ⊛ reflections

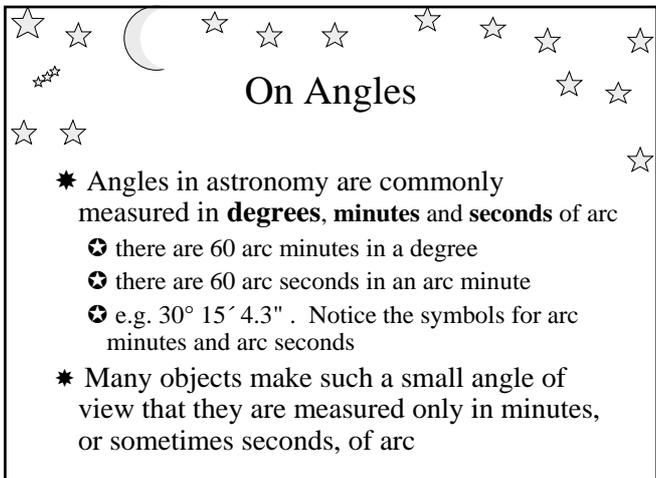

  
**On Vision**



- \* Small angles measure the ratio of *width/distance*
- \* The moon subtends an angle of about  $0.5^\circ$ 
  - ⊛ the moon is 0.27 times earth's diameter and at an average distance of 384000 km
- \* The sun is  $400 \times$  the moon's diameter and  $390 \times$  as far away. It, too, subtends an angle of about  $0.5^\circ$


  
**Angular Measurement**




  
**On Angles**

- \* Angles in astronomy are commonly measured in **degrees, minutes** and **seconds** of arc
  - ⊛ there are 60 arc minutes in a degree
  - ⊛ there are 60 arc seconds in an arc minute
  - ⊛ e.g.  $30^\circ 15' 4.3''$ . Notice the symbols for arc minutes and arc seconds
- \* Many objects make such a small angle of view that they are measured only in minutes, or sometimes seconds, of arc

**Example on Angles**

- \* Given: Saturn's diameter is 120,000 km  
Saturn's distance is 1500 million km
- ⊕ what angle does the disk of Saturn subtend?
- \* Calculate: 1. work out the ratio *width/distance*  
2. use *arctan* or *tan<sup>-1</sup>* calculator button  
3. answer is  $0.00458^\circ$   
( $\times 60$ )  $\equiv 0.275'$  (arc minutes)  
( $\times 60$ )  $\equiv 16.5''$  (arc seconds)

**The Eye**

- \* If your eye's lens is perfect, the size of the light receptor cells at the back of your eye sets the limit on how well you can distinguish close together objects



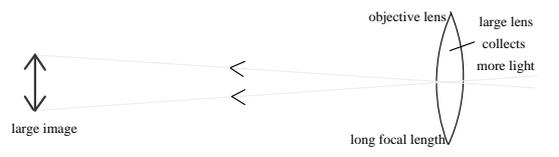
- \* The sharpest eyes can distinguish objects  $0.01^\circ$  apart, about  $0.5'$  arc
- ⊕ you could just tell whether someone is holding up 1 or 2 fingers at 100 m. You can't see Saturn's rings unaided

**The 8 Planets & More**

- \* Mercury
- \* Venus
- \* Earth
- \* Mars
- \* Jupiter      Asteroids
- \* Saturn
- \* Uranus - barely seen with naked eye
- \* Neptune - not seen with naked eye
- Pluto & more

Cue slides

**Purpose of a Telescope**

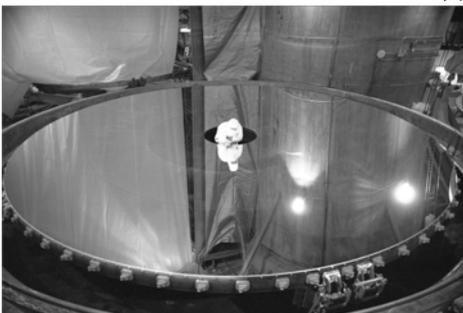


- \* The longer the focal length, the larger the image
- \* The bigger the lens, the more light collected
- ⊕ a telescope eyepiece gives additional magnification
- ⊕  $\times 400$  is a reasonable limit for normal viewing

**Modern telescopes**



- \* Above: a suite of international telescopes in Hawaii at an altitude of over 4000 m
- \* One of the Gemini 8 m diameter mirrors

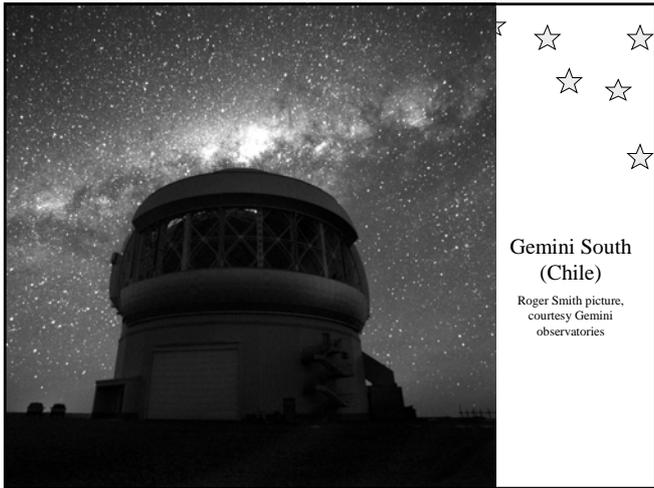


Images courtesy Gemini Observatory/AURA

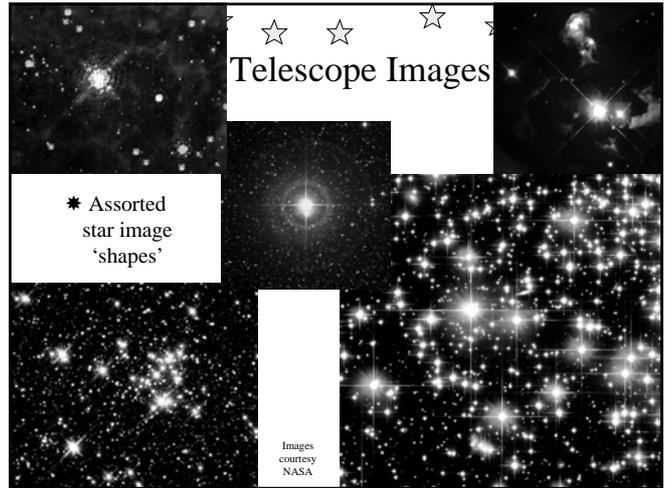


Moonset at Gemini North (Hawaii)

Courtesy: Gemini observatory/AURA



Gemini South  
(Chile)  
Roger Smith picture,  
courtesy Gemini  
observatories



### Telescope Images

\* Assorted  
star image  
'shapes'

Images  
courtesy  
NASA

### Units of Distance

- \* Earth's equatorial diameter is 12756 km
- \* **Astronomical Unit (AU)** is now defined as a fixed distance of about  $149.598 \times 10^6$  km, effectively the average distance between Earth and Sun, more than 10,000 Earth diameters.

- ☉ the AU is the natural unit to use for solar system distances
- ☉ Pluto's orbit is 80 AU in diameter, Jupiter's 10.4 AU
- \* Light takes about 8 minutes to travel 1 AU. In 1 year light travels  $9.5 \times 10^{12}$  km in space. This distance is called a **light year (LY)**.

### Scale of the Universe

- ☉ to Sun: 8 light min
- ☉ to nearest star: 4.2 LY
- ☉ across galaxy:  $10^5$  LY
- ☉ local cluster:  $\sim 10^7$  LY
- ☉ very distant galaxies:  $\sim 5 \times 10^9$  LY

Courtesy: K & K

animated

### Stars in our Galaxy

- \* There are over  $10^{11}$  stars in our galaxy - a truly astronomical number
- ☉ the number of grains of sand each occupying  $1 \text{ mm}^2$  that can be seen on a strip of Aberdeen's beach 1 km long by 100 m wide is:

Our sister galaxy  
Andromeda

← 1 km =  $10^6$  mm →

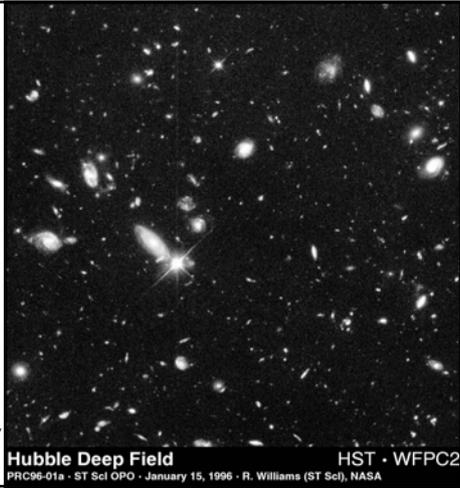
↑ 100 m =  $10^5$  mm ↓

$10^6 \times 10^5 = 10^{11}$  grains



☆ ☆ ☾  
☆☆  
**Hubble  
Deep Field**

- \* A stunning enlargement of our view of the universe
- \* Historical review



**Hubble Deep Field** HST · WFPC2  
PRC96-01a · ST ScI OPO · January 15, 1996 · R. Williams (ST ScI), NASA